

\vec{W} , Properties of a massive spin=1 particle

• Massive Maxwell Eqs:

$$\partial^\alpha \partial_\alpha W^\mu - \partial^\mu \partial_\alpha W^\alpha + m_w^2 W^\mu = 0$$

$$\text{Let } W^\mu(x) = \int d^4k e^{ik_\nu x^\nu} W^\mu(k)$$

$$\text{and } W^\mu(k) = \sum_i \epsilon_i^\mu(k) a(k) \quad \text{where } \epsilon_i \cdot \epsilon_j = -\delta_{ij}$$

$$\text{We have } [-k^2 g^{\mu\alpha} + k^\mu k^\alpha + m_w^2 g^{\mu\alpha}] W_\alpha(k) = 0$$

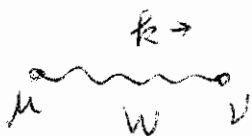
The propagator is the solution of

$$[-k^2 g^{\mu\alpha} + k^\mu k^\alpha + m_w^2 g^{\mu\alpha}] S_{\alpha\beta}(k) = i g_{\mu\beta}$$

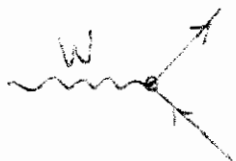
$$\Rightarrow S^{\alpha\beta}(k) = \frac{i}{k^2 - m_w^2} \left(-g^{\alpha\beta} + \frac{k^\alpha k^\beta}{m_w^2} \right)$$

↑ Same as $s=1$ ↑ Summation over polarizat

Feynman rules:



$$\frac{i}{k^2 - m_w^2} \left(-g^{\mu\nu} + \frac{k^\mu k^\nu}{m_w^2} \right)$$



$$i g_w \gamma^\mu (1 - \gamma_5)$$



$$i Q_f \gamma^\mu = -i e \gamma^\mu \quad (\text{for } e^-)$$

Polarization vectors of γ and the W 's ($S=1$)

- For γ , we have Maxwell's Eqs: ($m_\gamma = 0$, Spin = 1)

$$\nabla \cdot \vec{E} = \rho \quad (1) \quad \nabla \cdot \vec{B} = 0 \quad (2) \quad \nabla \times \vec{E} + \frac{\partial \vec{B}}{\partial t} = 0 \quad (3)$$

$$\nabla \times \vec{B} - \frac{\partial \vec{E}}{\partial t} = \vec{j} \quad (4)$$

- Use vector potential (\vec{A}, ϕ) where

$$\vec{E} = -\frac{\partial \vec{A}}{\partial t} - \nabla \phi \quad \text{and} \quad \vec{B} = \nabla \times \vec{A} \quad \text{Eqs (2) \& (3) are satisfied}$$

- Let $\mathcal{J}^\mu = (\rho, \vec{j})$, we have the covariant Maxwell Eqs.

$$\square^2 A^\mu - \partial^\mu \partial_\nu A^\nu = \mathcal{J}^\mu \quad \text{or} \quad \partial_\mu F^{\mu\nu} = \mathcal{J}^\nu, \quad \text{where } F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$$

- For $\mu=0$, we have Eq (1); and $\mu=1,3$ we have Eq (4).

- Gauge transformation $A'_\mu = A_\mu + \partial_\mu \Lambda(x)$

$$A_\mu \rightarrow A'_\mu = A_\mu + \partial_\mu \Lambda \quad \phi' = \phi - \frac{\partial \Lambda}{\partial t} \quad \text{Take } \Lambda = i a e^{-i \delta_\mu x}$$

- Define polarization vector $\xi_\mu(\vec{\delta})$ by $A_\mu = \xi_\mu(\vec{\delta}) e^{-i \delta_\nu x^\nu}$

$$\text{solution of } \square^2 A^\mu = 0$$

- Lorentz condition $\partial_\mu A^\mu = 0 \rightarrow \partial_\mu \xi^\mu = 0$

$$\text{Gauge Inv. } \xi'_\mu \Leftrightarrow \xi_\mu + a \delta_{\mu 0}$$

For $m_\gamma = 0$, take $t=0$, $\vec{\delta} \cdot \vec{\xi} = 0$ is transverse. For $\vec{P} = P_z \hat{z}$

$$\therefore \xi_1 = (0, 1, 0, 0), \quad \xi_2 = (0, 0, 1, 0) \quad \text{or}$$

$$\xi_{RL} = (\xi_1 \pm i \xi_2) / \sqrt{2}$$

- Spin = 0 particles obey Klein Gordon Eq: $i(\square^2 + m^2)\phi = -i\nabla\phi$

$$\square^2 \rightarrow -p^2 \quad \therefore \text{propagator} = \frac{1}{-p^2 + m^2}$$

Unitarity limit and the W^\pm mass limit
 We have

$$\sigma(\bar{\nu}_e + e^- \rightarrow \bar{\mu} + \bar{\nu}_\mu) = \frac{G^2 S}{3\pi}$$

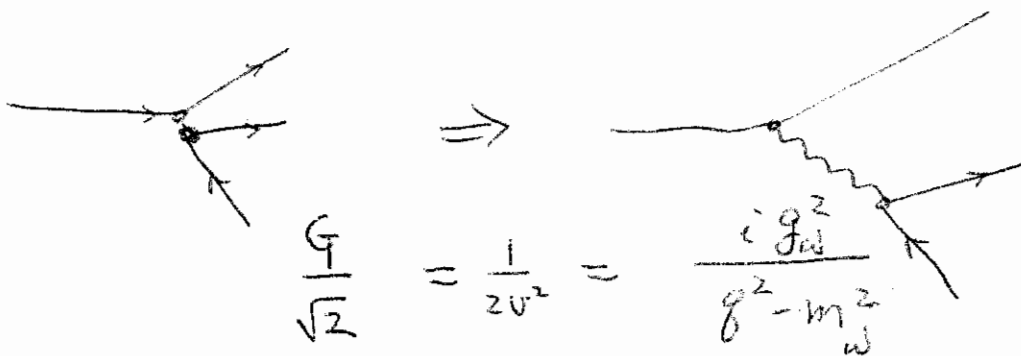
In the $J=1$ channel, $\bar{\nu}_e$ with RH only will contribute.
 (helicity)

$$\sigma(\bar{\nu}_e \text{ RH} + e^- \rightarrow \bar{\mu} + \bar{\nu}_\mu) \leq \frac{8\pi}{S} (2J+1) = \frac{24\pi}{S}$$

$$\therefore \frac{G^2 S}{3\pi} \leq \frac{24\pi}{S} \quad \therefore \sqrt{S}_{\text{violation}} = \left(\frac{6\pi\sqrt{2}}{G} \right)^{\frac{1}{2}} = 1516 \text{ GeV}$$

If we would save unitarity by using a form factor $G \rightarrow G(S)$
 which would mean leptons are not elementary or point-like

Introduction of W^\pm



$$\frac{G}{\sqrt{2}} = \frac{1}{2v^2} = \frac{i g_{\text{int}}^2}{g^2 - m_w^2} \quad v = (\sqrt{2} G)^{-\frac{1}{2}} = 246 \text{ GeV}$$

Consequences:

1) For $g^2 \ll m_w^2$, Fermi W.I. is recovered if $\frac{g_{\text{int}}^2}{m_w^2} = \frac{G}{\sqrt{2}}$

Perturbative calculation (which works well for W.I.) requires

From $e^+e^- \rightarrow W^+W^-$:
 $g_{\text{int}}^2 \ll 1$, i.e. $m_w \ll 350 \text{ GeV}$; $m_w = 80.37 \pm 0.16 \text{ GeV}$

Use similar argument, we'll set limit on m_t !!!

$$(2) \quad \sigma(\bar{\nu}_e + e^- \rightarrow \bar{\mu} + \bar{\nu}_\mu) = \frac{2g_w^4}{3\pi} \frac{s}{(s-m_w^2)^2}$$

$$= \frac{2g_w^4}{3\pi m_w^4} s = \frac{G_F^2 s}{3\pi} \quad \text{same as Fermi's W.I. as } s \ll m_w^2$$

As $s \rightarrow \infty \gg m_w^2$

$$\sigma = \frac{2g_w^4}{3\pi} \frac{1}{s} \leq \frac{24\pi}{s}$$

unitarity o.k. if $g_w^4 \leq 36\pi^2$

$$m_w < 1516 \text{ GeV}$$