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16.346 Astrodynamics
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Lagrange's Variational Methods for Linear Equations

Consider the equation

$$\frac{d^2y}{dt^2} + y = \sec t \quad \Longrightarrow \quad \begin{aligned} \frac{dy_1}{dt} &= y_2 \\ \frac{dy_2}{dt} + y_1 &= \sec t \end{aligned} \quad \Longrightarrow \quad \frac{d}{dt} \begin{bmatrix} y_1 \\ y_2 \end{bmatrix} = \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix} \begin{bmatrix} y_1 \\ y_2 \end{bmatrix} + \begin{bmatrix} 0 \\ \sec t \end{bmatrix}$$

which is equivalent to

$$\frac{d\mathbf{y}}{dt} = \mathbf{F}\mathbf{y} + \mathbf{g}$$

where

$$\mathbf{y} = \begin{bmatrix} y_1 \\ y_2 \end{bmatrix} \quad \text{and} \quad \mathbf{F} = \begin{bmatrix} 0 & 1 \\ -1 & 0 \end{bmatrix} \quad \text{and} \quad \mathbf{g} = \begin{bmatrix} 0 \\ \sec t \end{bmatrix}$$

Now the Wronskian matrix \mathbf{W}

$$\mathbf{W} = \begin{bmatrix} \sin t & -\cos t \\ \cos t & \sin t \end{bmatrix} \quad \text{satisfies} \quad \frac{d\mathbf{W}}{dt} = \mathbf{F}\mathbf{W}$$

and the solution of the homogeneous equation is

$$\mathbf{y}_h = \mathbf{W}\mathbf{c} \quad \text{where} \quad \mathbf{c} = \begin{bmatrix} c_1 \\ c_2 \end{bmatrix}$$

We now seek a solution of the general equation

$$\frac{d\mathbf{y}}{dt} = \mathbf{F}\mathbf{y} + \mathbf{g} \quad \text{of the form} \quad \mathbf{y} = \mathbf{W}\mathbf{c}(t)$$

Substitute and obtain

$$\frac{d\mathbf{W}}{dt}\mathbf{c} + \mathbf{W}\frac{d\mathbf{c}}{dt} = \mathbf{F}\mathbf{W}\mathbf{c} + \mathbf{g} \quad \text{which reduces to} \quad \mathbf{W}\frac{d\mathbf{c}}{dt} = \mathbf{g}$$

Hence

$$\frac{d\mathbf{c}}{dt} = \mathbf{W}^{-1}\mathbf{g}$$

which is solved by quadratures to obtain

$$\frac{d\mathbf{c}}{dt} = \begin{bmatrix} \sin t & \cos t \\ -\cos t & \sin t \end{bmatrix} \begin{bmatrix} 0 \\ \sec t \end{bmatrix} \quad \text{or} \quad \frac{dc_1}{dt} = 1 \quad \text{and} \quad \frac{dc_2}{dt} = \tan t$$

Hence

$$c_1(t) = t + c_1 \quad \text{and} \quad c_2(t) = \log(\sec t + \tan t) + c_2$$

so that the general solution is simply

$$\mathbf{y} = \mathbf{W}\mathbf{c}(t) \quad \text{or} \quad y(t) = c_1(t) \sin t - c_2(t) \cos t$$

Derivation of the Variational Equations

$$\frac{d}{dt} \begin{bmatrix} \mathbf{r} \\ \mathbf{v} \end{bmatrix} = \begin{bmatrix} \mathbf{O} & \mathbf{I} \\ -\frac{\mu}{r^3} \mathbf{I} & \mathbf{O} \end{bmatrix} \begin{bmatrix} \mathbf{r} \\ \mathbf{v} \end{bmatrix} + \begin{bmatrix} \mathbf{0} \\ \mathbf{a}_d \end{bmatrix} \iff \frac{d\mathbf{s}}{dt} = \mathbf{F} \mathbf{s} + \boldsymbol{\eta}$$

where

$$\mathbf{s} = \mathbf{s}(t, \boldsymbol{\alpha}) = \begin{bmatrix} \mathbf{r}(t, \boldsymbol{\alpha}) \\ \mathbf{v}(t, \boldsymbol{\alpha}) \end{bmatrix}$$

Two-Body Motion: $\frac{\partial \mathbf{s}}{\partial t} = \mathbf{F} \mathbf{s}$ **Disturbed Motion:** $\frac{d\mathbf{s}}{dt} = \mathbf{F} \mathbf{s} + \boldsymbol{\eta}$

Seek solutions of the form $\mathbf{s} = \mathbf{s}(t, \alpha_1, \alpha_2, \alpha_3, \alpha_4, \alpha_5, \alpha_6)$

where, for example, $\boldsymbol{\alpha}^T = [\Omega \quad i \quad \omega \quad a \quad e \quad \lambda = -n\tau]$ and $\boldsymbol{\eta} = \begin{bmatrix} \mathbf{0} \\ \mathbf{a}_d \end{bmatrix}$

Differentiate

$$\frac{d\mathbf{s}}{dt} = \frac{\partial \mathbf{s}}{\partial t} + \frac{\partial \mathbf{s}}{\partial \boldsymbol{\alpha}} \frac{d\boldsymbol{\alpha}}{dt} = \mathbf{F} \mathbf{s} + \boldsymbol{\eta} \implies \frac{\partial \mathbf{s}}{\partial \boldsymbol{\alpha}} \frac{d\boldsymbol{\alpha}}{dt} = \boldsymbol{\eta}$$

Since

$$\frac{\partial \mathbf{s}}{\partial \boldsymbol{\alpha}} \frac{\partial \boldsymbol{\alpha}}{\partial \mathbf{s}} = \frac{\partial \boldsymbol{\alpha}}{\partial \mathbf{s}} \frac{\partial \mathbf{s}}{\partial \boldsymbol{\alpha}} = \mathbf{I}$$

then

$$\frac{\partial \boldsymbol{\alpha}}{\partial \mathbf{s}} \frac{\partial \mathbf{s}}{\partial \boldsymbol{\alpha}} \frac{d\boldsymbol{\alpha}}{dt} = \frac{\partial \boldsymbol{\alpha}}{\partial \mathbf{s}} \boldsymbol{\eta} = \begin{bmatrix} \frac{\partial \boldsymbol{\alpha}}{\partial \mathbf{r}} & \frac{\partial \boldsymbol{\alpha}}{\partial \mathbf{v}} \end{bmatrix} \begin{bmatrix} \mathbf{0} \\ \mathbf{a}_d \end{bmatrix} = \frac{\partial \boldsymbol{\alpha}}{\partial \mathbf{v}} \mathbf{a}_d$$

so that

$$\boxed{\frac{d\boldsymbol{\alpha}}{dt} = \frac{\partial \boldsymbol{\alpha}}{\partial \mathbf{v}} \mathbf{a}_d}$$

$$\begin{aligned} x = f(\xi, \eta) \quad \begin{bmatrix} dx \\ dy \end{bmatrix} &= \begin{bmatrix} f_\xi & f_\eta \\ g_\xi & g_\eta \end{bmatrix} \begin{bmatrix} d\xi \\ d\eta \end{bmatrix} & \quad \xi = F(x, y) \quad \begin{bmatrix} d\xi \\ d\eta \end{bmatrix} &= \begin{bmatrix} F_x & F_y \\ G_x & G_y \end{bmatrix} \begin{bmatrix} dx \\ dy \end{bmatrix} \\ y = g(\xi, \eta) \end{aligned}$$

$$\begin{bmatrix} dx \\ dy \end{bmatrix} = \begin{bmatrix} f_\xi & f_\eta \\ g_\xi & g_\eta \end{bmatrix} \begin{bmatrix} F_x & F_y \\ G_x & G_y \end{bmatrix} \begin{bmatrix} dx \\ dy \end{bmatrix} \implies \begin{bmatrix} f_\xi & f_\eta \\ g_\xi & g_\eta \end{bmatrix} \begin{bmatrix} F_x & F_y \\ G_x & G_y \end{bmatrix} = \mathbf{I}$$

Variation of the Classical Elements

$$\mu \left(\frac{2}{r} - \frac{1}{a} \right) = v^2 = \mathbf{v} \cdot \mathbf{v} = \mathbf{v}^T \mathbf{v} \implies \frac{\mu}{a^2} \frac{\partial a}{\partial \mathbf{v}} = 2\mathbf{v}^T \implies \boxed{\frac{da}{dt} = \frac{2a^2}{\mu} \mathbf{v} \cdot \mathbf{a}_d}$$

$$\mathbf{h} = \mathbf{r} \times \mathbf{v} = \mathbf{S}_r \mathbf{v} \implies \frac{\partial \mathbf{h}}{\partial \mathbf{v}} = \mathbf{S}_r \frac{\partial \mathbf{v}}{\partial \mathbf{v}} = \mathbf{S}_r \mathbf{I} = \mathbf{S}_r \implies \boxed{\frac{d\mathbf{h}}{dt} = \mathbf{r} \times \mathbf{a}_d}$$

$$h^2 = \mathbf{h}^T \mathbf{h} \implies 2h \frac{\partial h}{\partial \mathbf{v}} = 2\mathbf{h}^T \mathbf{S}_r \implies \boxed{\frac{dh}{dt} = \mathbf{i}_h \cdot \mathbf{r} \times \mathbf{a}_d = \mathbf{i}_h \times \mathbf{r} \cdot \mathbf{a}_d = r \mathbf{i}_\theta \cdot \mathbf{a}_d}$$

or, alternately,

$$h^2 = (\mathbf{r} \times \mathbf{v}) \cdot (\mathbf{r} \times \mathbf{v}) = \mathbf{r}^T \mathbf{r} \mathbf{v}^T \mathbf{v} - \mathbf{r}^T \mathbf{v} \mathbf{r}^T \mathbf{v} \implies \boxed{\frac{dh}{dt} = \frac{1}{h} \mathbf{r}^T (\mathbf{r} \mathbf{v}^T - \mathbf{v} \mathbf{r}^T) \mathbf{a}_d}$$

$$p = \frac{h^2}{\mu} = a(1 - e^2) \implies \boxed{2\mu a e \frac{de}{dt} = \mu(1 - e^2) \frac{da}{dt} - 2h \frac{dh}{dt}}$$

$$\boxed{\mu \frac{d\mathbf{e}}{dt} = \mathbf{a}_d \times (\mathbf{r} \times \mathbf{v}) + (\mathbf{a}_d \times \mathbf{r}) \times \mathbf{v}}$$

Variation of i and Ω

From Page 84 in the textbook

$$\mathbf{h} = h \mathbf{i}_h = h(\sin \Omega \sin i \mathbf{i}_x - \cos \Omega \sin i \mathbf{i}_y + \cos i \mathbf{i}_z) \quad (2.6)$$

Then
$$\frac{d\mathbf{h}}{dt} = h \sin i \frac{d\Omega}{dt} \mathbf{i}_n - h \frac{di}{dt} \mathbf{i}_m + \frac{dh}{dt} \mathbf{i}_h$$

where

$$\mathbf{i}_n = \cos \Omega \mathbf{i}_x + \sin \Omega \mathbf{i}_y \quad (2.5)$$

$$\mathbf{i}_m = \mathbf{i}_h \times \mathbf{i}_n = -\sin \Omega \cos i \mathbf{i}_x + \cos \Omega \cos i \mathbf{i}_y + \sin i \mathbf{i}_z \quad (2.8)$$

Hence

$$\boxed{\begin{aligned} \frac{d\Omega}{dt} &= \frac{1}{h \sin i} \mathbf{i}_n \times \mathbf{r} \cdot \mathbf{a}_d = \frac{r \sin \theta}{h \sin i} \mathbf{i}_h \cdot \mathbf{a}_d \\ \frac{di}{dt} &= -\frac{1}{h} \mathbf{i}_m \times \mathbf{r} \cdot \mathbf{a}_d = \frac{r \cos \theta}{h} \mathbf{i}_h \cdot \mathbf{a}_d \end{aligned}}$$

where $\theta = \omega + f$ is the argument of latitude.

Variation of the true anomaly f

$$r(1 + e \cos f) = \frac{h^2}{\mu} \implies r e \sin f \frac{\partial f}{\partial \mathbf{v}} = r \cos f \frac{\partial e}{\partial \mathbf{v}} - \frac{2h}{\mu} \frac{\partial h}{\partial \mathbf{v}}$$

From Eq. (3.29) $\frac{\mu}{h} r e \sin f = \mathbf{r} \cdot \mathbf{v} \implies r e \cos f \frac{\partial f}{\partial \mathbf{v}} = -r \sin f \frac{\partial e}{\partial \mathbf{v}} + \frac{\mathbf{r} \cdot \mathbf{v}}{\mu} \frac{\partial h}{\partial \mathbf{v}} + \frac{h}{\mu} \mathbf{r}^T$

Multiply the first by $\cos f$, the second by $\sin f$ and add to obtain

$$r e h \frac{\partial f}{\partial \mathbf{v}} = (p \cos f) \mathbf{r}^T - (p + r) \sin f \frac{\partial h}{\partial \mathbf{v}} \quad \text{to be used in} \quad \boxed{\frac{df}{dt} = \frac{h}{r^2} + \frac{\partial f}{\partial \mathbf{v}} \mathbf{a}_d}$$

Variation of ω

$$\mathbf{i}_n = \cos \Omega \mathbf{i}_x + \sin \Omega \mathbf{i}_y \quad \Longrightarrow \quad \cos \theta = \mathbf{i}_n \cdot \mathbf{i}_r = \cos \Omega (\mathbf{i}_x \cdot \mathbf{i}_r) + \sin \Omega (\mathbf{i}_y \cdot \mathbf{i}_r)$$

Then

$$-\sin \theta \frac{\partial \theta}{\partial \mathbf{v}} = [-\sin \Omega (\mathbf{i}_x \cdot \mathbf{i}_r) + \cos \Omega (\mathbf{i}_y \cdot \mathbf{i}_r)] \frac{\partial \Omega}{\partial \mathbf{v}} \quad \Longrightarrow \quad \frac{\partial \theta}{\partial \mathbf{v}} = -\cos i \frac{\partial \Omega}{\partial \mathbf{v}}$$

since

$$\begin{aligned} \mathbf{i}_x \cdot \mathbf{i}_r &= \cos \Omega \cos \theta - \sin \Omega \sin \theta \cos i \\ \mathbf{i}_y \cdot \mathbf{i}_r &= \sin \Omega \cos \theta + \cos \Omega \sin \theta \cos i \end{aligned}$$

This gives the perturbative derivative of θ , i.e., the change in θ due to the change in \mathbf{i}_n from which the angle θ is measured. The total time rate of change of θ is the sum

$$\frac{d\theta}{dt} = \frac{\partial \theta}{\partial t} + \frac{\partial \theta}{\partial \mathbf{v}} \mathbf{a}_d = \frac{h}{r^2} - \cos i \frac{d\Omega}{dt}$$

Since $\theta = \omega + f$, then

$$\boxed{\frac{d\omega}{dt} = -\frac{\partial f}{\partial \mathbf{v}} \mathbf{a}_d - \cos i \frac{d\Omega}{dt}}$$

Gauss' form of Lagrange's variational equations in polar coordinates

$$\begin{aligned} \frac{d\Omega}{dt} &= \frac{r \sin \theta}{h \sin i} a_{dh} \\ \frac{di}{dt} &= \frac{r \cos \theta}{h} a_{dh} \\ \frac{d\omega}{dt} &= \frac{1}{he} [-p \cos f a_{dr} + (p+r) \sin f a_{d\theta}] - \frac{r \sin \theta \cos i}{h \sin i} a_{dh} \\ \frac{da}{dt} &= \frac{2a^2}{h} \left(e \sin f a_{dr} + \frac{p}{r} a_{d\theta} \right) \\ \frac{de}{dt} &= \frac{1}{h} \{ p \sin f a_{dr} + [(p+r) \cos f + re] a_{d\theta} \} \\ \frac{df}{dt} &= \frac{h}{r^2} + \frac{1}{eh} [p \cos f a_{dr} - (p+r) \sin f a_{d\theta}] \end{aligned}$$

Gauss' form of the variational equations in tangential-normal coordinates

$$\begin{aligned} \frac{d\omega}{dt} &= \frac{1}{ev} \left[2 \sin f a_{dt} + \left(2e + \frac{r}{a} \cos f \right) a_{dn} \right] - \frac{r \sin \theta \cos i}{h \sin i} a_{dh} \\ \frac{da}{dt} &= \frac{2a^2 v}{\mu} a_{dt} \\ \frac{de}{dt} &= \frac{1}{v} \left[2(e + \cos f) a_{dt} - \frac{r}{a} \sin f a_{dn} \right] \\ \frac{df}{dt} &= \frac{h}{r^2} - \frac{1}{ev} \left[2 \sin f a_{dt} + \left(2e + \frac{r}{a} \cos f \right) a_{dn} \right] \end{aligned}$$